Multipole fields

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1 Introduction

The big idea is to time- transform all physical quantities, such as the current vector and the fields. We have

$$J^{\mu}\left(\mathbf{\tilde{x}},t\right) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} J^{\mu}\left(\mathbf{\tilde{x}},\omega\right) \exp\left(-i\omega t\right) dt$$

The wave equation for A is:

$$\Box^2 \mathbf{A} = \frac{4\pi}{c} \mathbf{J}$$

and we already have the solution

$$\mathbf{A} = \frac{4\pi}{c} \int \mathbf{J} \left(\mathbf{x}' \right) \frac{\delta \left(\left| \mathbf{\tilde{x}} - \mathbf{\tilde{x}}' \right| - c \left(t - t' \right) \right)}{4\pi \left| \mathbf{\tilde{x}} - \mathbf{\tilde{x}}' \right|} d^4 \mathbf{x}'$$

which we may write in terms of the time transform of \mathbf{J}

$$A^{\mu}(\mathbf{x}) = \frac{1}{c} \int \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} J^{\mu}(\tilde{\mathbf{x}}', \omega) e^{-i\omega t'} d\omega \frac{\delta\left(\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right| - c\left(t - t'\right)\right)}{\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|} d^{4}\mathbf{x}'$$

Now we do the integration over t':

$$A^{\mu}(\mathbf{x}) = \frac{1}{c} \int \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} J^{\mu}(\mathbf{\tilde{x}}',\omega) e^{-i\omega t'} d\omega \frac{\delta\left(ct' - c(t - |\mathbf{\tilde{x}} - \mathbf{\tilde{x}}'|/c)\right)}{|\mathbf{\tilde{x}} - \mathbf{\tilde{x}}'|} dct' d^{3}\mathbf{\tilde{x}}'$$
$$= \frac{1}{\sqrt{2\pi}} \int \int_{-\infty}^{+\infty} \frac{J^{\mu}(\mathbf{\tilde{x}}',\omega) \exp\left(-i\omega(t - |\mathbf{\tilde{x}} - \mathbf{\tilde{x}}'|/c)\right)}{c|\mathbf{\tilde{x}} - \mathbf{\tilde{x}}'|} d\omega d^{3}\mathbf{\tilde{x}}'$$
(1)

The 0th component is:

$$\Phi\left(\mathbf{x}\right) = \frac{1}{\sqrt{2\pi}} \int \int_{-\infty}^{+\infty} d\omega e^{-i\omega t} \frac{\rho\left(\tilde{\mathbf{x}}',\omega\right) \exp\left(i\omega\left|\tilde{\mathbf{x}}-\tilde{\mathbf{x}}'\right|/c\right)}{\left|\tilde{\mathbf{x}}-\tilde{\mathbf{x}}'\right|} d^{3}\tilde{\mathbf{x}}'$$
(2)

while the 1,2,3 components are:

$$\tilde{\mathbf{A}} = \frac{1}{\sqrt{2\pi}} \int \int_{-\infty}^{+\infty} d\omega e^{-i\omega t} \frac{\tilde{\mathbf{J}}\left(\tilde{\mathbf{x}}',\omega\right) \exp\left(i\omega \left|\tilde{\mathbf{x}}-\tilde{\mathbf{x}}'\right|/c\right)}{c \left|\tilde{\mathbf{x}}-\tilde{\mathbf{x}}'\right|} d^{3}\tilde{\mathbf{x}}'$$
(3)

Transforming the relations between the fields and the potentials, we get:

$$\tilde{\mathbf{B}}\left(\tilde{\mathbf{x}},\omega\right) = \tilde{\boldsymbol{\nabla}} \times \tilde{\mathbf{A}}\left(\tilde{\mathbf{x}},\omega\right)$$

and

$$\frac{1}{c}\frac{\partial \mathbf{E}}{\partial t} = \tilde{\boldsymbol{\nabla}} \times \tilde{\mathbf{B}} \to -\frac{i\omega}{c}\tilde{\mathbf{E}}\left(\tilde{\mathbf{x}},\omega\right) = \tilde{\boldsymbol{\nabla}} \times \tilde{\mathbf{B}}\left(\tilde{\mathbf{x}},\omega\right)$$

and thus

$$\tilde{\mathbf{E}}\left(\tilde{\mathbf{x}},\omega\right) = \frac{ic}{\omega}\tilde{\boldsymbol{\nabla}}\times\tilde{\mathbf{B}}\left(\tilde{\mathbf{x}},\omega\right)$$

We have 3 relevant length scales: d, the dimension of the source, λ , the wavelength, and r, the distance from source to observer. The ordering of these lengths determines how we proceed.

- d ≪ r ≪ λ. This is the near, or static region. With r ≪ λ, the exponential in equations(3) and (2) is exp(2πir/λ) ≈ 1 and we get the static results from Chapter 3 for the time transform à (x̃, ω). Thus we get the static fields, but oscillating in time.
- $d \ll r \sim \lambda$. The induction zone. This is tricky.
- d ≪ λ ≪ r. The radiation zone. In this zone the source appears almost point-like. We may expand the quantity |x̃ − x̃'|:

$$\begin{aligned} \left| \mathbf{\tilde{x}} - \mathbf{\tilde{x}}' \right|^2 &= r^2 + x'^2 - 2\mathbf{\tilde{r}} \cdot \mathbf{\tilde{x}}' \\ \left| \mathbf{\tilde{x}} - \mathbf{\tilde{x}}' \right| &= r \left(1 - \frac{\mathbf{\hat{r}} \cdot \mathbf{\tilde{x}}'}{r} \right) = r - \mathbf{\hat{r}} \cdot \mathbf{\tilde{x}}' \end{aligned}$$

and similarly

$$\frac{1}{\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|} = \frac{1}{r} + \hat{\mathbf{r}} \cdot \tilde{\mathbf{x}}' + \dots \approx \frac{1}{r}$$

Remember that we need more accuracy in the exponential than in the quantity outside the exponential. From here on we shall assume we are in the radiation zone.

2 The dipole fields

We put the approximations for $|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'|$ into the expression for $\tilde{\mathbf{A}}$ (equation 3):

$$\tilde{\mathbf{A}} = \frac{1}{\sqrt{2\pi}} \int \int_{-\infty}^{+\infty} d\omega e^{-i\omega t} \frac{\tilde{\mathbf{J}}\left(\tilde{\mathbf{x}}',\omega\right) \exp\left(ikr - ik\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right)}{cr} d^{3}\tilde{\mathbf{x}}'$$

Thus the time transform of $\tilde{\mathbf{A}}$ is:

$$\tilde{\mathbf{A}}\left(\tilde{\mathbf{x}},\omega\right) = \frac{e^{ikr}}{cr} \int \tilde{\mathbf{J}}\left(\tilde{\mathbf{x}}',\omega\right) \exp\left(-ik\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right) d^{3}\tilde{\mathbf{x}}'$$

where $k = \omega/c$. Now we expand the exponential in the integrand:

$$\exp\left(-ik\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right) = 1 - ik\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}' + \frac{1}{2}\left(-ik\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right)^2 + \cdots$$
$$= 1 - ik\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}' - \frac{k^2}{2}\left(\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right)^2 + \cdots$$

The first term in $\mathbf{\tilde{A}}$ is:

$$\tilde{\mathbf{A}}_{d}\left(\tilde{\mathbf{x}},\omega\right) = \frac{e^{ikr}}{cr} \int \tilde{\mathbf{J}}\left(\tilde{\mathbf{x}}',\omega\right) d^{3}\tilde{\mathbf{x}}'$$

We can simplify this expression by using the equation of charge conservation:

$$\frac{\partial \rho}{\partial t} + \mathbf{\tilde{\nabla}} \cdot \mathbf{\tilde{J}} = 0$$

Taking the time transform, we have:

$$-i\omega
ho+\mathbf{ ilde{\nabla}}\cdot\mathbf{ ilde{J}}\left(\mathbf{ ilde{x}},\omega
ight)=0$$

Now

$$\begin{split} \tilde{\boldsymbol{\nabla}}\left(\tilde{\mathbf{x}}\cdot\tilde{\mathbf{J}}\right) &= \left(\tilde{\mathbf{x}}\cdot\tilde{\boldsymbol{\nabla}}\right)\tilde{\mathbf{J}} + \left(\tilde{\mathbf{J}}\cdot\tilde{\boldsymbol{\nabla}}\right)\tilde{\mathbf{x}} + \tilde{\mathbf{x}}\times\left(\tilde{\boldsymbol{\nabla}}\times\tilde{\mathbf{J}}\right) + \tilde{\mathbf{J}}\times\left(\tilde{\boldsymbol{\nabla}}\times\tilde{\mathbf{x}}\right) \\ &= \left(\tilde{\mathbf{x}}\cdot\tilde{\boldsymbol{\nabla}}\right)\tilde{\mathbf{J}} + \tilde{\mathbf{J}} + \tilde{\mathbf{x}}\times\left(\tilde{\boldsymbol{\nabla}}\times\tilde{\mathbf{J}}\right) \\ &= x_{m}\partial_{m}J_{i} + J_{i} + \varepsilon_{ijk}x_{j}\varepsilon_{klm}\partial_{l}J_{m} \\ &= x_{m}\partial_{m}J_{i} + J_{i} + \left(\delta_{il}\delta_{jm} - \delta_{im}\delta_{jl}\right)x_{j}\partial_{l}J_{m} \\ &= x_{m}\partial_{m}J_{i} + J_{i} + x_{j}\partial_{i}J_{j} - x_{l}\partial_{l}J_{i} \\ &= J_{i} + x_{j}\partial_{i}J_{j} + x_{i}\partial_{m}J_{m} - x_{i}\partial_{m}J_{m} \\ &= \tilde{\mathbf{J}} + \tilde{\mathbf{x}}\left(\tilde{\boldsymbol{\nabla}}\cdot\tilde{\mathbf{J}}\right) + x_{j}\partial_{i}J_{j} - x_{i}\partial_{m}J_{m} \\ &= \tilde{\mathbf{J}} + \tilde{\mathbf{x}}\left(\tilde{\boldsymbol{\nabla}}\cdot\tilde{\mathbf{J}}\right) + \left(\tilde{\mathbf{x}}\times\tilde{\boldsymbol{\nabla}}\right)\times\tilde{\mathbf{J}} \end{split}$$
(4)

and

$$\begin{aligned} \int_{\text{all space}} \tilde{\boldsymbol{\nabla}} \left(\tilde{\mathbf{x}} \cdot \tilde{\mathbf{J}} \right) dV &= \int_{s_{\infty}} \left(\tilde{\mathbf{x}} \cdot \tilde{\mathbf{J}} \right) \hat{\mathbf{n}} dA = 0 \\ &= \int_{\text{all space}} \left(\tilde{\mathbf{J}} + \tilde{\mathbf{x}} \left(\tilde{\boldsymbol{\nabla}} \cdot \tilde{\mathbf{J}} \right) + \left(\tilde{\mathbf{x}} \times \tilde{\boldsymbol{\nabla}} \right) \times \tilde{\mathbf{J}} \right) dV \end{aligned}$$

since \mathbf{J} is zero outside the source. Thus:

$$\int \tilde{\mathbf{J}} \left(\tilde{\mathbf{x}}', \omega \right) d^3 \tilde{\mathbf{x}}' = -\int \tilde{\mathbf{x}}' \left(\tilde{\boldsymbol{\nabla}}' \cdot \tilde{\mathbf{J}} \right) d^3 \tilde{\mathbf{x}}'$$
$$= -i \int \tilde{\mathbf{x}}' \omega \rho \left(\tilde{\mathbf{x}}', \omega \right) d^3 \tilde{\mathbf{x}}' = -i \omega \tilde{\mathbf{p}} \left(\omega \right)$$

where $\mathbf{\tilde{p}}$ is the dipole moment of the source.

Then:

$$\tilde{\mathbf{A}}_{d}\left(\tilde{\mathbf{x}},\omega\right) = -i\omega\frac{e^{ikr}}{cr}\tilde{\mathbf{p}}\left(\omega\right) = -ik\frac{e^{ikr}}{r}\tilde{\mathbf{p}}\left(\omega\right)$$
$$\tilde{\mathbf{B}} = \tilde{\boldsymbol{\nabla}} \times \tilde{\mathbf{A}} = -\frac{ik}{r}\left(ik - \frac{1}{r}\right)e^{ikr}\hat{\mathbf{r}} \times \tilde{\mathbf{p}}$$

and in the radiation zone $k\gg 1/r,$ so:

$$\tilde{\mathbf{B}} = k^2 \frac{e^{ikr}}{r} \hat{\mathbf{r}} \times \tilde{\mathbf{p}}$$

and then

$$\begin{split} \tilde{\mathbf{E}} \left(\tilde{\mathbf{x}}, \omega \right) &= \frac{i}{k} \tilde{\boldsymbol{\nabla}} \times \tilde{\mathbf{B}} \left(\tilde{\mathbf{x}}, \omega \right) = \frac{i}{k} \tilde{\boldsymbol{\nabla}} \times \left(k^2 \frac{e^{ikr}}{r} \hat{\mathbf{r}} \times \tilde{\mathbf{p}} \right) \\ &= ik \left(ik \frac{e^{ikr}}{r} \right) \hat{\mathbf{r}} \times \left(\hat{\mathbf{r}} \times \tilde{\mathbf{p}} \right) \\ &= -k^2 \frac{e^{ikr}}{r} \hat{\mathbf{r}} \times \left(\hat{\mathbf{r}} \times \tilde{\mathbf{p}} \right) \end{split}$$

3 Power radiated

The power radiated per unit solid angle is given by the Poynting vector:

$$\frac{dP}{d\Omega} = r^2 \left| \mathbf{\tilde{S}} \right| = r^2 \frac{c}{4\pi} \left| \mathbf{\tilde{E}} \times \mathbf{\tilde{B}} \right|$$

The the total energy radiated is

$$\frac{dW}{d\Omega} = \frac{c}{4\pi} \int_{-\infty}^{+\infty} r^2 \left| \mathbf{\tilde{E}} \times \mathbf{\tilde{B}} \right| dt$$

and using Parseval's theorem we may convert to an integral of the transforms over frequency:

$$\frac{dW}{d\Omega} = \frac{c}{4\pi} \int_{-\infty}^{+\infty} r^2 \left| \mathbf{\tilde{E}} \times \mathbf{\tilde{B}}^* \right| d\omega$$

where

$$r^{2} \left| \tilde{\mathbf{E}} \times \tilde{\mathbf{B}}^{*} \right| = r^{2} \frac{c}{4\pi} \left| \left(-k^{2} \frac{e^{ikr}}{r} \hat{\mathbf{r}} \times (\hat{\mathbf{r}} \times \tilde{\mathbf{p}}) \right) \times \left(k^{2} \frac{e^{-ikr}}{r} \hat{\mathbf{r}} \times \tilde{\mathbf{p}}^{*} \right) \right|$$
$$= \frac{c}{4\pi} k^{4} \left| \hat{\mathbf{r}} \times \tilde{\mathbf{p}} \right|^{2}$$

and so

$$\frac{d^2 W}{d\Omega d\omega} = \frac{c}{4\pi} k^4 \left| \hat{\mathbf{r}} \times \tilde{\mathbf{p}} \right|^2 \tag{5}$$

is the energy radiated per unit solid angle and per unit frequency.

4 Periodic source

If the source is periodic, then the current must be expanded in a Fourier series rather than a Fourier transform. We have

$$\mathbf{J}\left(\mathbf{\tilde{x}},t\right) = \sum_{n=-\infty}^{+\infty} \mathbf{J}_{n}\left(\mathbf{\tilde{x}}\right) \exp\left(in\omega_{0}t\right)$$

where ω_0 is the fundamental frequency = $2\pi/T$ and T is the period of the source. Then the

expression for $\mathbf{\tilde{A}}$ becomes:

$$\mathbf{A}(\mathbf{x}) = \frac{1}{c} \int \sum_{n=-\infty}^{+\infty} \mathbf{J}_n(\tilde{\mathbf{x}}') \exp(in\omega_0 t') \frac{\delta\left(\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right| - c\left(t - t'\right)\right)}{\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|} d^4 \mathbf{x}'$$
$$= \sum_{n=-\infty}^{+\infty} \int \frac{\mathbf{J}_n(\tilde{\mathbf{x}}) \exp(in\omega_0 t)}{c\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|} \exp\left(-in\omega_0\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|/c\right) d^3 \tilde{\mathbf{x}}'$$

which is a Fourier series for A with coefficients:

$$\mathbf{A}_{n} = \sum_{n=-\infty}^{+\infty} \int \frac{\mathbf{J}_{n}\left(\tilde{\mathbf{x}}\right)}{c \left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|} \exp\left(-in\omega_{0}\left|\tilde{\mathbf{x}} - \tilde{\mathbf{x}}'\right|/c\right) d^{3}\tilde{\mathbf{x}}'$$

Following the analysis above, we find the dipole term to be:

$$\tilde{\mathbf{A}}_{n,d}\left(\tilde{\mathbf{x}}\right) = -in\omega_0 \frac{e^{-ik_n r}}{cr} \tilde{\mathbf{p}}_n = -ink_0 \frac{e^{-ik_n r}}{r} \tilde{\mathbf{p}}_n$$

where $k_n = n\omega_0/c$ and

$$\tilde{\mathbf{p}}_{n} = \int \tilde{\mathbf{x}}' \rho_{n} \left(\tilde{\mathbf{x}}' \right) d^{3} \tilde{\mathbf{x}}'$$

with ρ_n being the *n*th coefficient in the Fourier series for ρ . Then we find the power radiated is:

$$\begin{aligned} \frac{dP}{d\Omega} &= r^2 \frac{c}{4\pi} \left| \mathbf{\tilde{E}} \times \mathbf{\tilde{B}} \right| \\ &= r^2 \frac{c}{4\pi} \left| \left(\sum_{n=-\infty}^{+\infty} -k_n^2 \frac{e^{-ik_n r}}{r} \mathbf{\hat{r}} \times (\mathbf{\hat{r}} \times \mathbf{\tilde{p}}_n) e^{in\omega_o t} \right) \times \left(\sum_{m=-\infty}^{+\infty} k_m^2 \frac{e^{-ik_m r}}{r} \mathbf{\hat{r}} \times \mathbf{\tilde{p}}_m e^{im\omega_0 t} \right) \right| \\ &= \frac{c}{4\pi} \sum_{n=-\infty}^{+\infty} \sum_{m=-\infty}^{+\infty} k_n^2 k_m^2 e^{-ik_n r} e^{-ik_m r} \left| -(\mathbf{\hat{r}} \times (\mathbf{\hat{r}} \times \mathbf{\tilde{p}}_n)) \times (\mathbf{\hat{r}} \times \mathbf{\tilde{p}}_m) \right| \exp\left(i\omega_0 t (n+m)\right) \end{aligned}$$

Now we time average by integrating over one period T and dividing by T. Then only those terms with m = -n survive the integration. We have:

$$<\frac{dP}{d\Omega}>=\frac{c}{4\pi}\sum_{n=-\infty}^{+\infty}k_{n}^{4}\left|\hat{\mathbf{r}}\times\tilde{\mathbf{p}}_{n}\right|^{2}$$
(6)

This equation is very similar to equation (5), but the meaning is somewhat different. Equation (5) is the energy radiated per unit frequency – the power spectrum. On the other hand equation (6) gives the time averaged power in the *n*th harmonic for a periodic source. Note however, that the real physical frequency ω_n is described by both the negative and the positive frequency $\pm \omega_n$, and for a real dipole $\tilde{\mathbf{p}}_n^* = \tilde{\mathbf{p}}_{-n}$, so

$$<\frac{dP_n}{d\Omega}>=\frac{c}{2\pi}k_n^4\left|\hat{\mathbf{r}}\times\tilde{\mathbf{p}}_n\right|^2\tag{7}$$

In this expression $\bar{\mathbf{p}}_n$ is the coefficient of the *exponential* Fourier series. Careful use of these expressions obviates any of the factor-of-2 problems which otherwise arise.

Example: Suppose we have a pure frequency dipole: $\mathbf{\tilde{p}} = \mathbf{\tilde{p}}_0 \cos(\omega_0 t) =$

$$\begin{split} \frac{\tilde{\mathbf{p}}_0}{2} \left(e^{i\omega_0 t} + e^{-i\omega_0 t} \right). \ \text{Thus } \tilde{\mathbf{p}}_1 &= \tilde{\mathbf{p}}_{-1} = \tilde{\mathbf{p}}_0/2. \text{ Then} \\ \tilde{\mathbf{A}} &= -ink_0 \frac{1}{2r} \tilde{\mathbf{p}}_0 \left(e^{ik_0 r} e^{i\omega_0 t} - e^{-ik_0 r} e^{-i\omega_0 t} \right) \end{split}$$

and

$$\begin{aligned} \frac{dP}{d\Omega} &= r^2 \frac{c}{4\pi} \left| \tilde{\mathbf{E}} \times \tilde{\mathbf{B}} \right| \\ &= r^2 \frac{c}{4\pi} \left| \left(-k_0^2 \frac{1}{2r} \hat{\mathbf{r}} \times (\hat{\mathbf{r}} \times \tilde{\mathbf{p}}_0) \left(e^{ik_0 r} e^{i\omega_o t} - e^{-ik_0 r - i\omega_0 t} \right) \right) \times \left(k_0^2 \frac{1}{2r} \hat{\mathbf{r}} \times \tilde{\mathbf{p}}_0 \left(e^{ik_0 r} e^{i\omega_o t} - e^{-ik_0 r - i\omega_0 t} \right) \right) \\ &= \frac{k_0^4}{4} \frac{c}{4\pi} \left| \hat{\mathbf{r}} \times \tilde{\mathbf{p}}_0 \right|^2 \left| 2 - e^{2(ik_0 r)} e^{i2\omega_o t} - e^{-2(ik_0 r)} e^{-i2\omega_o t} \right| \end{aligned}$$

Now we time average to get:

$$<\frac{dP}{d\Omega}>=2\frac{c}{16\pi}k_{0}^{4}\left|\mathbf{\hat{r}}\times\mathbf{\tilde{p}}_{0}\right|^{2}=\frac{c}{8\pi}k_{0}^{4}\left|\mathbf{\hat{r}}\times\mathbf{\tilde{p}}_{0}\right|^{2}$$

which agrees with Jackson's equation 9.22.

Alternatively, we can use equation (7) directly. We get:

$$<\frac{dP_n}{d\Omega}>=\frac{c}{2\pi}k_n^4\left|\hat{\mathbf{r}}\times\tilde{\mathbf{p}}_n\right|^2=\frac{c}{2\pi}\left(\frac{\omega_0}{c}\right)^4\left|\hat{\mathbf{r}}\times\frac{\tilde{\mathbf{p}}_0}{2}\right|^2=\frac{c}{8\pi}\left(\frac{\omega_0}{c}\right)^4\left|\hat{\mathbf{r}}\times\tilde{\mathbf{p}}_0\right|^2$$

as expected.

5 Quadrupole fields

The second term in the expansion of $\tilde{\mathbf{A}}$ is:

$$\tilde{\mathbf{A}}_{2} = \frac{e^{ikr}}{cr} \int \tilde{\mathbf{J}} \left(\tilde{\mathbf{x}}', \omega \right) \left(-ik\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}}' \right) d^{3}\tilde{\mathbf{x}}'$$

To evaluate this term, first we work on the inetgrand. Note that

$$\left(\mathbf{ ilde{x}}' imes \mathbf{ ilde{J}}
ight) imes \mathbf{ ilde{r}} = \mathbf{ ilde{J}} \left(\mathbf{ ilde{r}} \cdot \mathbf{ ilde{x}}'
ight) - \mathbf{ ilde{x}}' \left(\mathbf{ ilde{r}} \cdot \mathbf{ ilde{J}}
ight)$$

Thus we can write

$$\frac{1}{c}\tilde{\mathbf{J}}\left(\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right) = \frac{1}{2c}\left\{\tilde{\mathbf{J}}\left(\hat{\mathbf{r}}\cdot\tilde{\mathbf{x}}'\right) + \tilde{\mathbf{x}}'\left(\hat{\mathbf{r}}\cdot\tilde{\mathbf{J}}\right) + \left(\tilde{\mathbf{x}}'\times\tilde{\mathbf{J}}\right)\times\hat{\mathbf{r}}\right\}$$

The antisymmetric part of this expression is

$$\frac{1}{2c} \left(\mathbf{\tilde{x}}' \times \mathbf{\tilde{J}} \right) \times \mathbf{\hat{r}} = \mathbf{\tilde{M}} \times \mathbf{\hat{r}}$$

where $\tilde{\mathbf{M}}$ is the magnetization (cf Jackson equation 5.53). Thus:

$$\begin{aligned} \tilde{\mathbf{A}}_{2} &= -ik\frac{e^{ikr}}{r} \left(\int \frac{1}{2c} \left\{ \tilde{\mathbf{J}} \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}}' \right) + \tilde{\mathbf{x}}' \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{J}} \right) \right\} d^{3}\tilde{\mathbf{x}}' + \tilde{\mathbf{m}} \times \hat{\mathbf{r}} \right) \\ &= \tilde{\mathbf{A}}_{q} + \tilde{\mathbf{A}}_{md} \end{aligned}$$

where

$$\tilde{\mathbf{A}}_{md} = -ik\frac{e^{ikr}}{r}\tilde{\mathbf{m}} \times \hat{\mathbf{r}}$$

is the magnetic dipole term and

$$\tilde{\mathbf{A}}_{q} = -ik\frac{e^{ikr}}{r} \int \frac{1}{2c} \left\{ \tilde{\mathbf{J}} \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}}' \right) + \tilde{\mathbf{x}}' \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{J}} \right) \right\} d^{3}\tilde{\mathbf{x}}'$$

is the quadrupole term.

Analysis of the magnetic dipole term proceeds exactly as for the electric dipole term, with $\tilde{\mathbf{p}}$ replaced by $\tilde{\mathbf{m}} \times \hat{\mathbf{r}}$. The extra cross product with $\hat{\mathbf{r}}$ changes the directions of $\tilde{\mathbf{E}}$ and $\tilde{\mathbf{B}}$, i.e. the polarization of these fields differs from that of the electric dipole terms.

$$\tilde{\mathbf{B}}_{md} = k^2 \frac{e^{ikr}}{r} \hat{\mathbf{r}} \times (\tilde{\mathbf{m}} \times \hat{\mathbf{r}}) = k^2 \frac{e^{ikr}}{r} \left(\tilde{\mathbf{m}} - \hat{\mathbf{r}} \left(\tilde{\mathbf{m}} \cdot \hat{\mathbf{r}} \right) \right)$$

while

$$\tilde{\mathbf{E}}_{md} = -k^2 \frac{e^{ikr}}{r} \left(\hat{\mathbf{r}} \times \tilde{\mathbf{m}} \right)$$

and the power is given by equation (4) or (6) with $\tilde{\mathbf{p}}$ replaced by $\tilde{\mathbf{m}}$.

Now we proceed with the quadrupole term. We are supposed to do an integration by parts, so let's run through the usual steps. First note that:

$$\partial_i \left(x_k J_i \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) \right) = \delta_{ik} J_i \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) + x_k \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) \tilde{\boldsymbol{\nabla}} \cdot \tilde{\mathbf{J}} + x_k J_i \left(\hat{\mathbf{r}} \cdot \tilde{\boldsymbol{\nabla}} \right) \tilde{\mathbf{x}}$$
$$= J_k \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) + x_k \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) \tilde{\boldsymbol{\nabla}} \cdot \tilde{\mathbf{J}} + x_k \tilde{\mathbf{J}} \cdot \hat{\mathbf{r}}$$

So

$$\int \partial_i \left[x_k J_i \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) \right] dV = \int_S \left(x_k J_i \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) \right) dA_i = 0$$
$$= \int \left(J_k \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) + x_k \left(\hat{\mathbf{r}} \cdot \tilde{\mathbf{x}} \right) \tilde{\boldsymbol{\nabla}} \cdot \tilde{\mathbf{J}} + x_k \tilde{\mathbf{J}} \cdot \hat{\mathbf{r}} \right) dV$$

Then:

$$\begin{split} \vec{A}_q &= -ik\frac{e^{ikr}}{r} \int \frac{1}{2c} \left\{ \vec{J} \left(\hat{\mathbf{r}} \cdot \vec{x}' \right) + \vec{x}' \left(\hat{\mathbf{r}} \cdot \vec{J} \right) \right\} d^3 \vec{x}' \\ &= ik\frac{e^{ikr}}{2cr} \int \vec{x} \left(\hat{\mathbf{r}} \cdot \vec{x} \right) \left(\vec{\nabla} \cdot \vec{J} \right) dV \\ &= ik\frac{e^{ikr}}{2cr} \int \vec{x} \left(\hat{\mathbf{r}} \cdot \vec{x} \right) i\omega \rho dV \\ &= -k^2 \frac{e^{ikr}}{2r} \int \vec{x} \left(\hat{\mathbf{r}} \cdot \vec{x} \right) \rho dV \end{split}$$

Then

$$\vec{B} = \nabla \times \vec{A} = -\frac{ik^3}{2r}e^{ikr} \int \left(\hat{\mathbf{r}} \times \vec{x}\right) \left(\hat{\mathbf{r}} \cdot \vec{x}\right) \rho dV$$

Now define the vector $\vec{Q}(\hat{\mathbf{r}})$ by

$$Q_i = \sum_j Q_{ij} \hat{r}_j \tag{8}$$

,

where Q_{ij} is the quadrupole tensor

$$Q_{ij} = \int \left(3x_i x_j - |\mathbf{\tilde{x}}|^2 \,\delta_{ij} \right) \rho\left(\mathbf{\tilde{x}}\right) d^3 \mathbf{\tilde{x}}$$

Then

$$Q_{i} = \sum_{j} \int \left(3x_{i}x_{j}\hat{r}_{j} - |\mathbf{\tilde{x}}|^{2}\hat{r}_{i} \right) \rho(\mathbf{\tilde{x}}) d^{3}\mathbf{\tilde{x}}$$
$$\vec{Q} = \int 3\mathbf{\tilde{x}} \left(\mathbf{\hat{r}} \cdot \mathbf{\tilde{x}} \right) \rho dV - \mathbf{\hat{r}} \int |\mathbf{\tilde{x}}|^{2} \rho dV$$

and thus

$$\vec{B} = -\frac{ik^3}{2r}e^{ikr}\frac{\hat{\mathbf{r}}\times\vec{Q}}{3}$$

and then

$$\vec{E} = \frac{i}{k} \left(\frac{k^4}{2r} e^{ikr} \frac{\hat{\mathbf{r}} \times \left(\hat{\mathbf{r}} \times \vec{Q} \right)}{3} \right)$$
$$= \frac{ik^3}{6r} e^{ikr} \hat{\mathbf{r}} \times \left(\hat{\mathbf{r}} \times \vec{Q} \right)$$

and thus

$$\frac{d^2 W}{d\Omega d\omega} = \frac{c}{4\pi} \frac{k^6}{36} \left| \hat{\mathbf{r}} \times \vec{Q} \right|^2$$
$$= \frac{c}{144\pi} k^6 \left| \hat{\mathbf{r}} \times \vec{Q} \right|^2$$

or, for a periodic source, the time averaged power is

$$<\frac{dP_n}{d\Omega}>=\frac{c}{72\pi}k^6\left|\hat{\mathbf{r}}\times\vec{Q_n}\right|^2$$

where again Q_n is the coefficient in the exponential Fourier series.

6 Angular distribution

For the dipole:

$$\frac{dP}{d\Omega} \propto \left| \hat{\mathbf{r}} \times \vec{p} \right|^2 = p^2 \sin^2 \theta$$

where θ is the angle between \vec{p} and the direction of propagation $\hat{\mathbf{r}}$.



Dipole radiation pattern

The total power radiated is:

$$P_{n} = \int \frac{dP_{n}}{d\Omega} d\Omega = \frac{c}{2\pi} k_{n}^{4} p_{n}^{2} (2\pi) \int_{-1}^{+1} (1-\mu^{2}) d\mu$$
$$= \frac{4}{3} c k_{n}^{4} p_{n}^{2}$$

For the quadrupole,

$$\frac{dP}{d\Omega} \propto \left| \hat{\mathbf{r}} \times \vec{Q}_n \left(\hat{\mathbf{r}} \right) \right|^2$$

Suppose a charge q oscillates along the z-axis from z = 0 to z = a, with period T. The dipole moment is:

$$p_z = zq = q\frac{a}{2}(1 + \cos\omega t) = q\frac{a}{4}(2 + e^{i\omega t} + e^{-i\omega t})$$

and the quadrupole moments are

which can be quite complicated.

$$Q_{11} = Q_{22} = -z^2 q = -q \frac{a^2}{4} (1 + \cos \omega t)^2$$

= $-q \frac{a^2}{4} (1 + 2\cos \omega t + \cos^2 \omega t)$
= $-q \frac{a^2}{4} \left(1 + 2\cos \omega t + \frac{\cos 2\omega t + 1}{2}\right)$
= $-q \frac{a^2}{4} \left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right)$

$$Q_{33} = 2z^2q = q\frac{a^2}{2}\left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right)$$

The dipole power radiated is

$$\frac{dP_n}{d\Omega} = \frac{c}{2\pi} k_n^4 p_n^2$$

and is all at the fundamental $\omega = \frac{2\pi}{T}$. We have $p_1 = qa/4$, so

$$\frac{dP_1}{d\Omega} = \frac{c}{2\pi} \left(\frac{\omega}{c}\right)^4 \left(\frac{qa}{4}\right)^2 \sin^2\theta = \frac{q^2 a^2 \omega^4}{32\pi c^3} \sin^2\theta$$

For the quadrupole, we first evaluate the vector $\vec{Q}\left(\hat{\mathbf{r}}
ight)$

$$Q_i = \sum_j Q_{ij} r_j$$

Thus

$$Q_1 = Q_{11}\sin\theta\cos\phi = -q\frac{a^2}{4}\left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right)\sin\theta\cos\phi$$
$$Q_2 = Q_{22}\sin\theta\sin\phi = -q\frac{a^2}{4}\left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right)\sin\theta\sin\phi$$
and
$$Q_3 = Q_{33}\cos\theta = q\frac{a^2}{2}\left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right)\cos\theta$$

Then

$$\begin{aligned} \hat{\mathbf{r}} \times \vec{Q} &= (\sin\theta\cos\phi, \sin\theta\sin\phi, \cos\theta) \times (\sin\theta\cos\phi, \sin\theta\sin\phi, -2\cos\theta) Q_{11} \\ &= (3\sin\theta\cos\theta\sin\phi, -3\cos\theta\sin\theta\cos\phi, 0) q \frac{a^2}{4} \left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right) \\ &= \left(\frac{3}{2}\sin 2\theta\sin\phi, -\frac{3}{2}\sin 2\theta\cos\phi, 0\right) q \frac{a^2}{4} \left(\frac{3}{2} + e^{i\omega t} + e^{-i\omega t} + \frac{e^{i2\omega t} + e^{-i2\omega t}}{4}\right) \end{aligned}$$

This vector has components at both ω and 2ω .

$$\begin{aligned} \frac{dP_1}{d\Omega} &= \frac{c}{72\pi} k^6 \left| \hat{\mathbf{r}} \times \vec{Q}_1\left(\hat{\mathbf{r}} \right) \right|^2 \\ &= \frac{c}{72\pi} k^6 \left(q \frac{a^2}{4} \right)^2 \left(\frac{9}{4} \sin^2 2\theta \right) \\ &= \frac{\omega^6 q^2 a^4}{512 c^5 \pi} \sin^2 2\theta \end{aligned}$$

while at the second harmonic:

$$\frac{dP_2}{d\Omega} = \frac{c}{72\pi} k^6 \left(q\frac{a^2}{16}\right)^2 \left(\frac{9}{4}\sin^2 2\theta\right)$$

$$= \frac{c}{72\pi} \left(\frac{2\omega}{c}\right)^6 \left(q\frac{a^2}{16}\right)^2 \left(\frac{9}{4}\sin^2 2\theta\right)$$

$$= \frac{\omega^6 q^2 a^4}{128c^5\pi} \sin^2 2\theta$$
(9)

The distribution given in equation (9) is shown below:



The total power radiated at ω is the sum of the dipole and quadrupole terms:

$$\frac{dP_1}{d\Omega} = \frac{q^2 a^2 \omega^4}{32\pi c^3} \sin^2 \theta + \frac{\omega^6 q^2 a^4}{2048\pi c^5} \sin^2 2\theta = \frac{q^2 a^2 \omega^4}{2048 c^3 \pi} \left(64 \sin^2 \theta + \frac{\omega^2 a^2}{c^2} \sin^2 2\theta \right)$$

The total power radiated in the quadrupole is:

$$P_{1} = \int \frac{dP}{d\Omega} d\Omega = \frac{c}{72\pi} k^{6} \int \left| \hat{\mathbf{r}} \times \vec{Q}_{n} \left(\hat{\mathbf{r}} \right) \right|^{2} d\Omega$$

where

$$\left(\mathbf{\hat{r}}\times\vec{Q}_{n}\left(\mathbf{\hat{r}}\right)\right)_{i}=\varepsilon_{ijk}r_{j}Q_{kl}r_{l}$$

and thus

$$\begin{aligned} \left| \hat{\mathbf{r}} \times \vec{Q}_{n} \left(\hat{\mathbf{r}} \right) \right|^{2} &= \varepsilon_{ijk} r_{j} Q_{kl} r_{l} \varepsilon_{imn} r_{m} Q_{np}^{*} r_{p} \\ &= \left(\delta_{jm} \delta_{kn} - \delta_{jn} \delta_{km} \right) r_{j} Q_{kl} r_{l} r_{m} Q_{np}^{*} r_{p} \\ &= r_{m} r_{m} Q_{kl} Q_{kp}^{*} r_{l} r_{p} - r_{j} r_{p} Q_{ml} Q_{jp}^{*} r_{l} r_{m} \\ &= r_{l} r_{p} \left(Q_{kl} Q_{kp}^{*} - r_{j} r_{m} Q_{ml} Q_{jp}^{*} \right) \\ &= \left. \vec{Q} \cdot \vec{Q}^{*} - \left| \hat{\mathbf{r}} \cdot \vec{Q} \right|^{2} \end{aligned}$$

The angle integrals give:

$$\int r_l r_p \delta\Omega = \begin{cases} 0 & \text{if } l \text{ or } p = 1, 2 \text{ and } l \neq p \\ \frac{4}{3}\pi & \text{if } l \text{ or } p = 1, 2 \text{ and } l = p \\ 2\pi \frac{2}{3} & \text{if } l = p = 3 \end{cases} = \frac{4\pi}{3} \delta_{lp}$$

while for

$$\int r_l r_p r_j r_m d\Omega$$

we note that if any of the indices l, p, j, m = 1 or 2, then one more of them must also equal that same index to survive the integration over ϕ . We need the results

$$\int_{0}^{2\pi} \cos \phi d\phi = \int_{0}^{2\pi} \sin \phi d\phi = 0$$
$$\int_{0}^{2\pi} \cos^{2} \phi d\phi = \int_{0}^{2\pi} \sin^{2} \phi d\phi = \pi$$
$$\int_{0}^{2\pi} \cos^{3} \phi d\phi = \int_{0}^{2\pi} \sin^{3} \phi d\phi = 0$$
$$\int_{0}^{2\pi} \cos^{4} \phi d\phi = \int_{0}^{2\pi} \sin^{4} \phi d\phi = \frac{3}{4}\pi$$
$$\int_{0}^{2\pi} \sin^{2} \phi \cos^{2} \phi d\phi = \frac{\pi}{4}$$

Each of the components of $\hat{\mathbf{r}}$ also contains either $\sin \theta$ or $\cos \theta$, and since we need an even number of $\sin \phi$ and $\cos \phi$ terms, we never have an odd power of $\sin \theta$, so we need:

$$\int_{-1}^{+1} \mu^4 d\mu = \frac{2}{5}$$

$$\int_{-1}^{+1} \mu^2 (1-\mu^2) d\mu = \frac{\mu^3}{3} - \frac{\mu^5}{5} \Big|_{-1}^{+1} = \frac{4}{15}$$

$$\int_{-1}^{+1} (1-\mu^2)^2 d\mu = \frac{16}{15}$$

These couple as follows:

$$\int r_l r_p r_j r_m d\Omega = \begin{cases} \pi \frac{4}{15} & \text{if } 2 \text{ of indices } = 1, 2 \text{ and } 2 \text{ indices } = 3\\ 2\pi \frac{2}{5} & \text{if } all \text{ indices } = 3\\ \frac{\pi}{4} \frac{16}{15} & \text{if } 2 \text{ indices } = 1 \text{ and } 2 \text{ indices } = 2\\ \frac{3\pi}{4} \frac{16}{15} & \text{if } all \text{ indices } = 1 \text{ or } 2. \end{cases}$$

Thus we may write:

$$\int r_l r_p r_j r_m d\Omega = \frac{4\pi}{15} \left(\delta_{lp} \delta_{jm} + \delta_{lj} \delta_{pm} + \delta_{lm} \delta_{jp} \right)$$

So we have

$$< P_{n} > = \frac{c}{72\pi} k^{6} \int r_{l} r_{p} \left(Q_{kl} Q_{kp}^{*} - r_{j} r_{m} Q_{ml} Q_{jp}^{*} \right) d\Omega$$

$$= \frac{c}{72\pi} k^{6} \left[Q_{kl} Q_{kp}^{*} \frac{4\pi}{3} \delta_{lp} - Q_{ml} Q_{jp}^{*} \frac{4\pi}{15} \left(\delta_{lp} \delta_{jm} + \delta_{lj} \delta_{pm} + \delta_{lm} \delta_{jp} \right) \right]$$

$$= \frac{c}{72\pi} k^{6} \frac{4\pi}{3} \left[Q_{kp} Q_{kp}^{*} - \frac{Q_{mp} Q_{mp}^{*}}{5} - \frac{Q_{mj} Q_{jm}^{*}}{5} - \frac{Q_{mm} Q_{pp}^{*}}{5} \right]$$

But Q_{ij} is traceless, so the last term is zero, so

$$< P_{n} > = \frac{\omega_{n}^{6}}{54c^{5}} \frac{3}{5} Q_{kp} Q_{kp}^{*}$$
$$= \frac{\omega_{n}^{6}}{90c^{5}} \sum_{k,p} |Q_{kp}|^{2}$$
(10)

Let's recalculate the power radiated by our oscillating charge. The quadrupole term at 2ω gives

$$< P_2 > = \frac{(2\omega)^6}{90c^5} \left(q\frac{a^2}{16}\right)^2 [1+1+4]$$
$$= \frac{1}{60} \frac{\omega^6}{c^5} q^2 a^4$$

Now compare with the integral of our previous term (9)

$$< P_{2} > = \int \frac{\omega^{6} q^{2} a^{4}}{128 c^{5} \pi} \sin^{2} 2\theta d\Omega$$

$$= \frac{\omega^{6} q^{2} a^{4}}{128 c^{5} \pi} (2\pi) \int_{-1}^{+1} 4 (1 - \mu^{2}) \mu^{2} d\mu$$

$$= \frac{\omega^{6} q^{2} a^{4}}{128 c^{5} \pi} (2\pi) (4) \frac{4}{15}$$

$$= \frac{1}{60} \frac{\omega^{6}}{c^{5}} q^{2} a^{4}$$

while at the fundamental

$$< P_1 > = \int \frac{\omega^6 q^2 a^4}{512 c^5 \pi} \sin^2 2\theta d\Omega = \frac{\omega^6 q^2 a^4}{512 c^5 \pi} \frac{32\pi}{15} = \frac{1}{240} \frac{\omega^6}{c^5} q^2 a^4$$

and the total power formula gives:

$$< P_1 > = \frac{\omega^6}{90c^5} \left(q\frac{a^2}{4}\right)^2 (1+1+4) = \frac{1}{240} \frac{\omega^6}{c^5} q^2 a^4$$

These expressions agree.